Experimental dark matter search at China Jinping underground Laboratory

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Shanghai Jiao Tong University

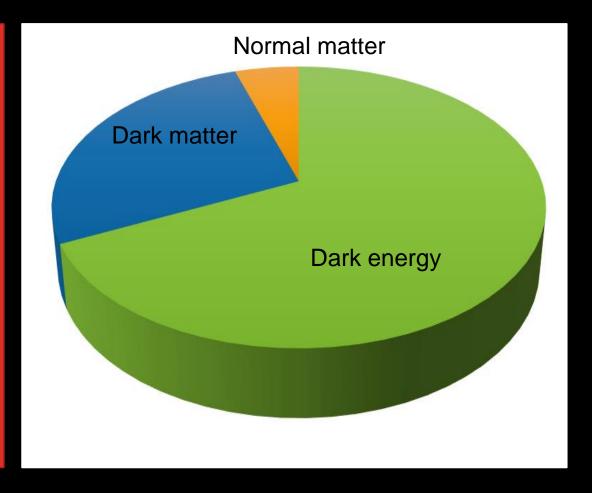


The GR view of the Universe

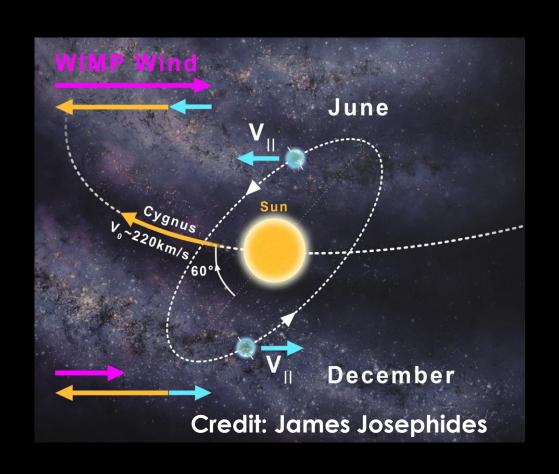


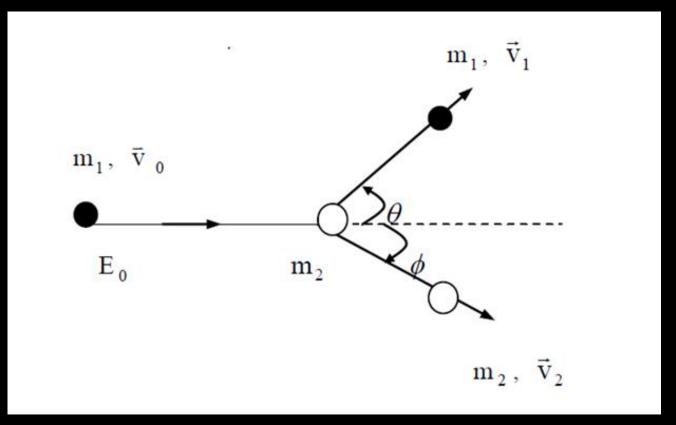
Law of universe expansion

Dark Energy All matter and energy in the universe

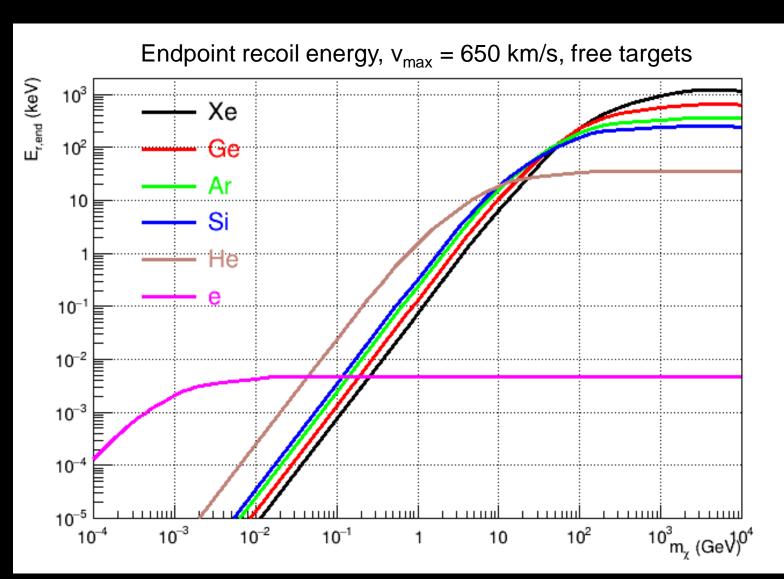


DM direct detection: classical beam experiment

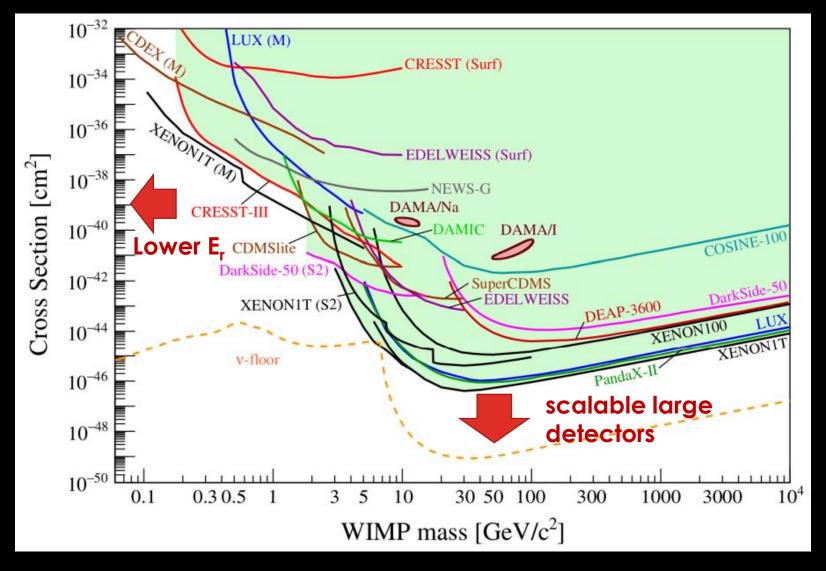




Visible recoil energy



Global status: hide & seek

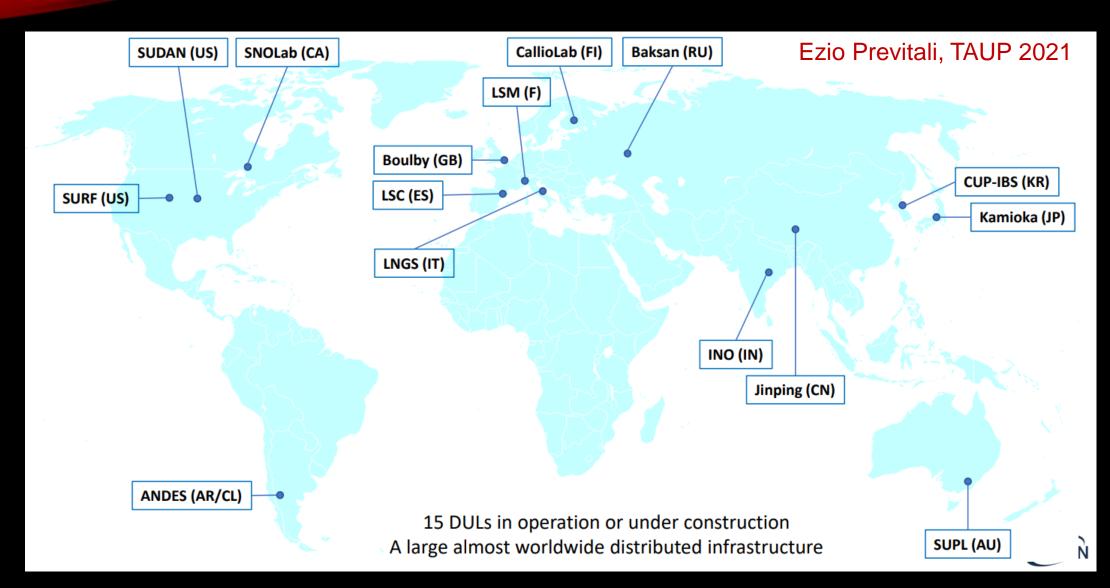


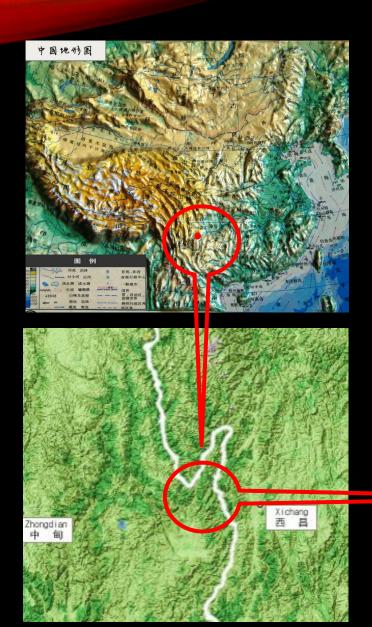
Name of the game: background

- <1 nucleus in our body is hit <u>per</u><u>year</u>!
- But our body is hit 10⁸/day by environmental background radiation!
- Hide detector in deep underground lab, and put massive shield

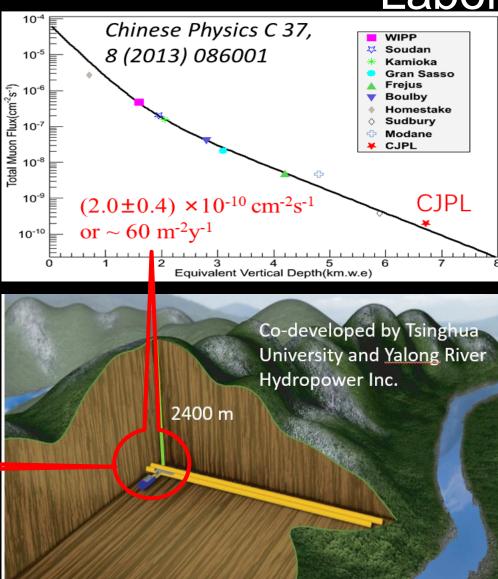


Underground labs



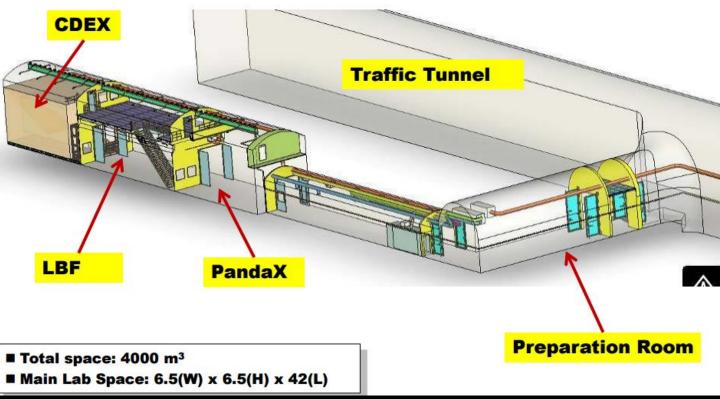


China Jinping underground <u>Laboratory</u>



CJPL-I

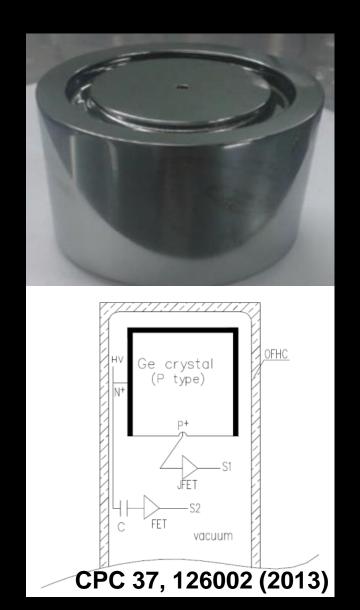




CDEX Collaboration

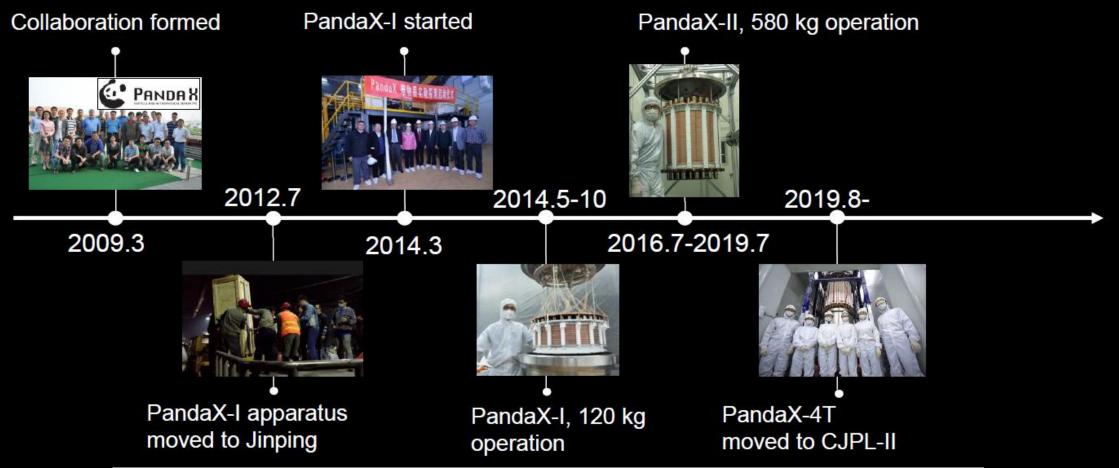
- China Dark Matter EXperiment
 - ✓ Formed in 2009, 11 institutions and ~70 people;
 - ✓ Searching for light DM by P-type Point-Contact Germanium detectors







Particle and Astrophysical Xenon observatory

























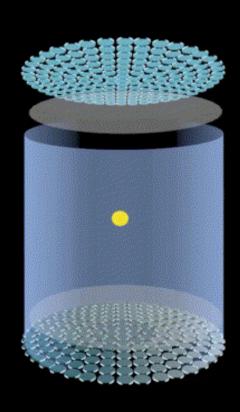




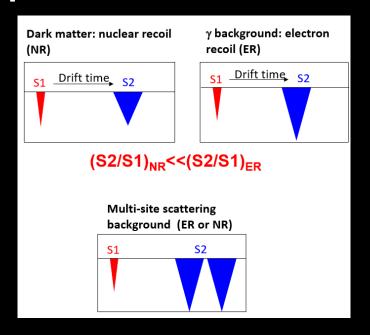




Gamma electron recoil WIMP nuclear recoil



Dual phase xenon TPC



Detector capability:

- Large monolithic target
- 3D reconstruction and fiducialization
- Good ER/NR rejection
- Calorimeter capable of seeing a couple of photons/electrons

CJPL-II Project (2014-)



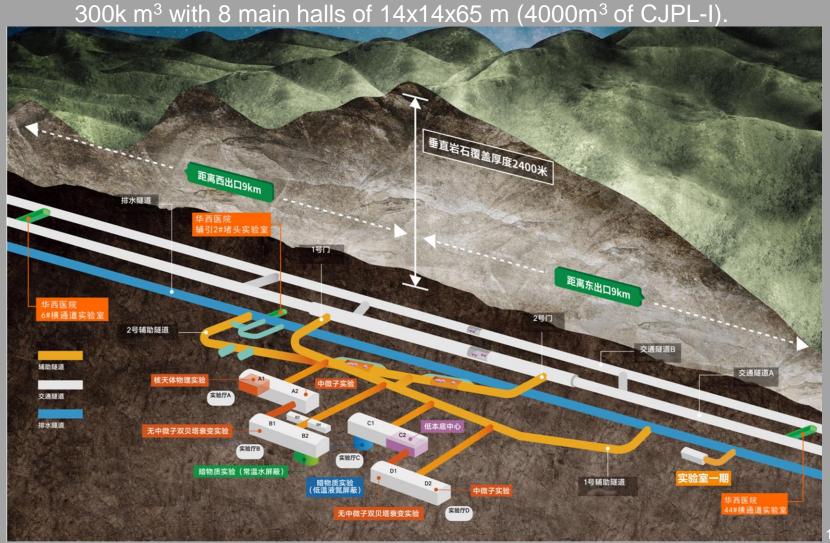
PHYSICS

China supersizes its underground physics lab

Planned expansion could pave way for "ultimate dark matter experiment"

he world's deepest physics laboratory is about to become one of its largest. Early next year, workers will start carving four cavernous experiment halls along a tunnel through Jinping Mountain in China's Sichuan province. Once the science at the China Jinping Underground Laboratory (CJPL) is scaled up as well, "it will be a milestone for Chinese physics," says Nigel Smith, director of the underground SNOLAB in Sudbury, Canada.

sics Science, Nov. 30, 2014



CJPL-II Project (2014-)



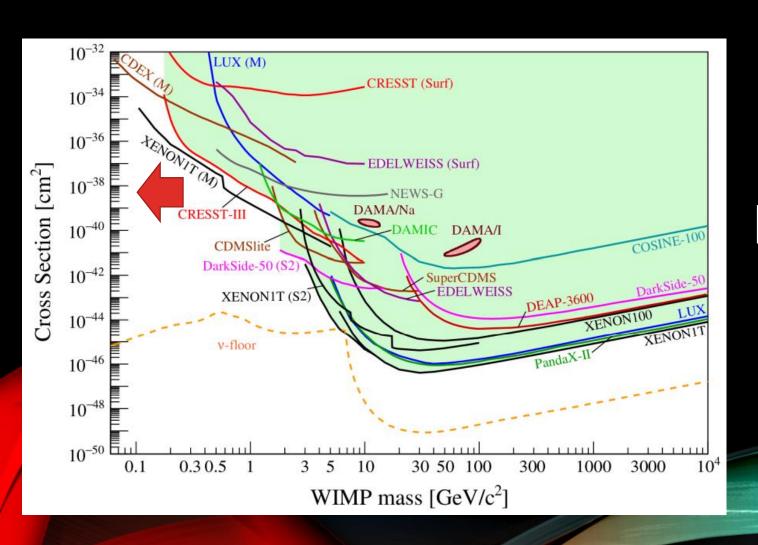




DURF @ CJPL-II

- December 2016, the National Development and Reform Commission of China issued the National Major Science and Technology Infrastructure Construction Projects for the 13th Five-year Plan
- DURF @ CJPL-II: Deep Underground and ultra-low Radiation background Facility for frontier physics experiments
- July 2019, DURF project commencement
- Second half, 2022, electromechanical equipment installation
- Early 2025, formal operation of DURF

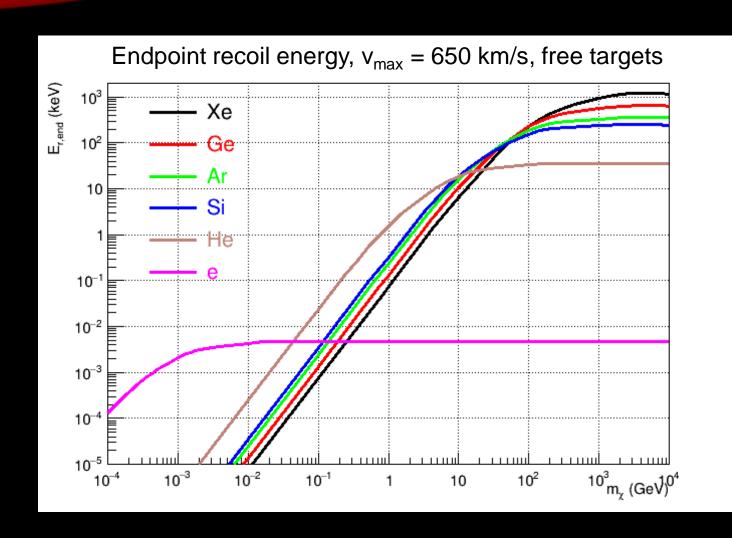




Light DM searches

Sub-GeV or so

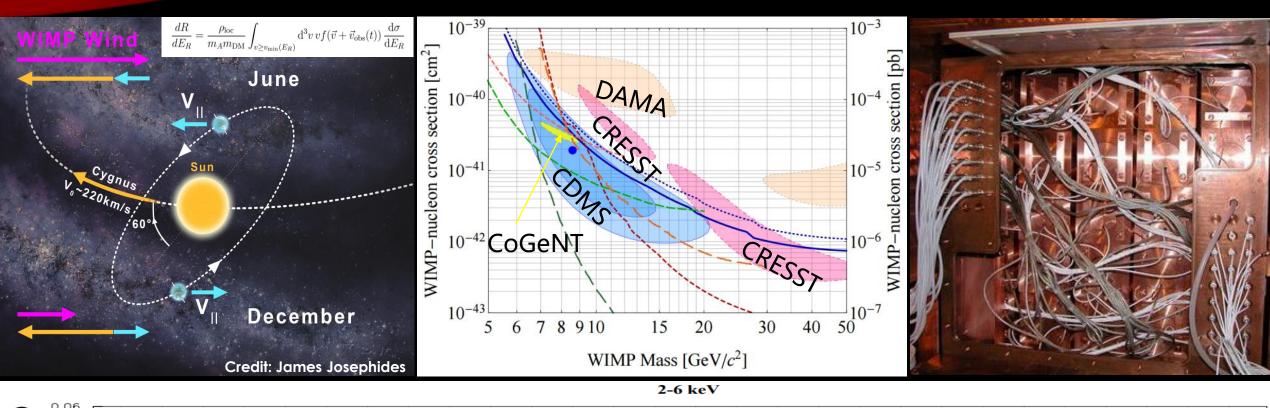
Probing low mass DM

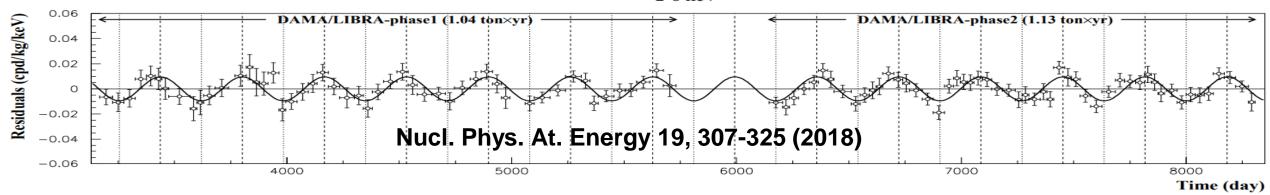


Strategies

- 1. Lowering energy threshold
- 2. Going after electrons (via atomic effects)
- 3. Boosting

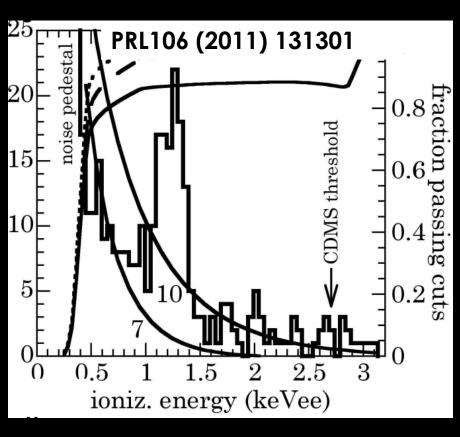
Light DM Anomaly

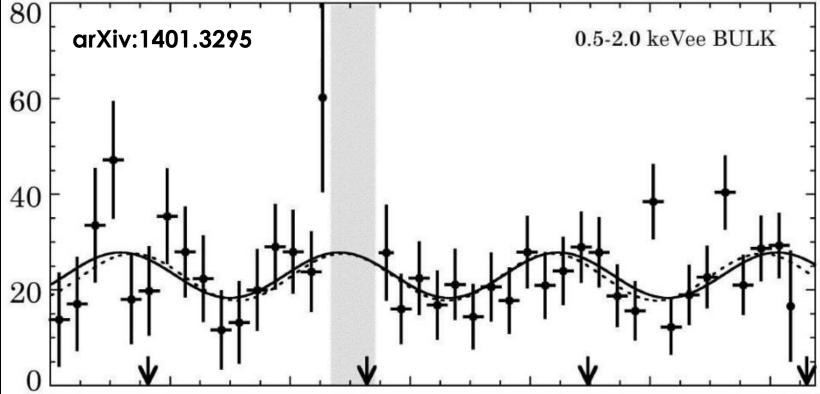




CoGeNT Anomaly

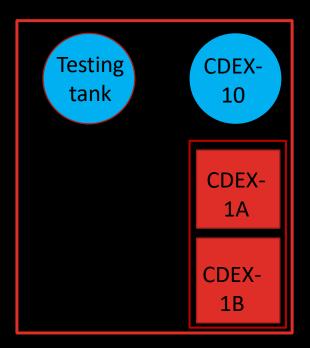
CoGeNT: P-type point contact Ge detector

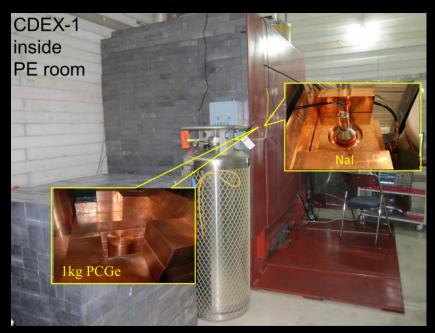


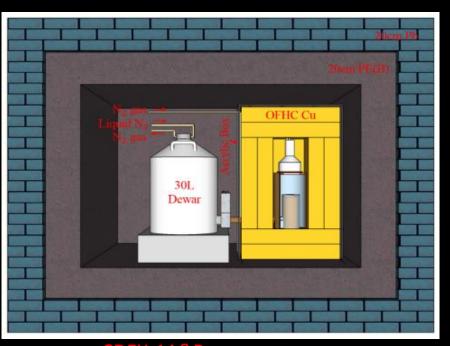


CDEX-1

- 2 sub-stages: CDEX-1A(prototype, 2011)→1B(upgraded, 2013);
- Single-element ~1kg PPC Ge detector w/ cold finger;
- Low-bkg Pb&Cu passive shield + Nal veto detector;
- Located in PE room at CJPL-I



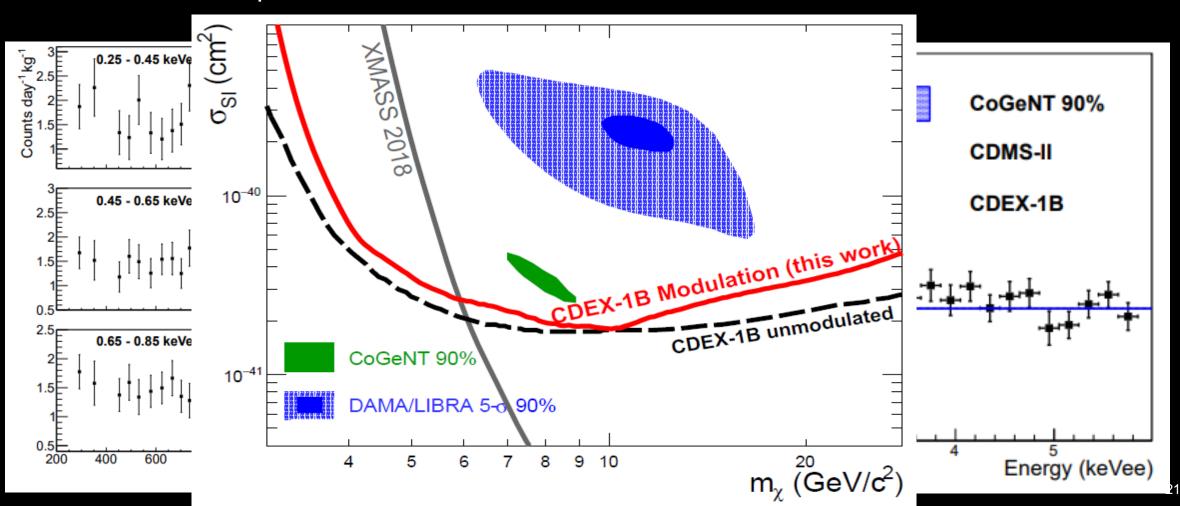




CDEX-1A&B

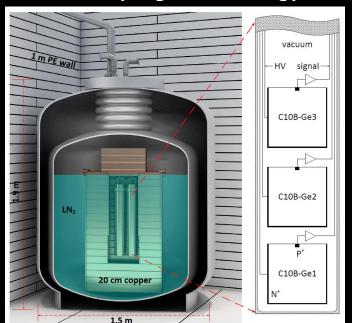
CDEX-1B results

 Annual Modulation analysis: more than 4 years of data, excluding CoGeNT and DAMA/LIBRA interpretation



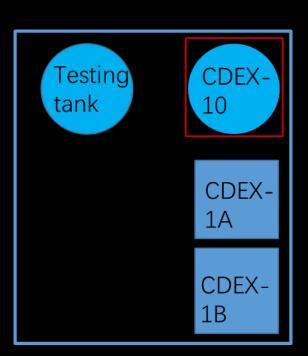
CDEX-10 Status

- Array detectors: 3 strings with 3 detectors each, ~10kg total;
- Direct immersion in LN₂;
- Prototype system for future hundred-kg to ton scale experiment
 - Light/radio-purer LN₂ replacing heavy shield i.e. Pb/Cu;
 - Arraying technology to scalable capability;



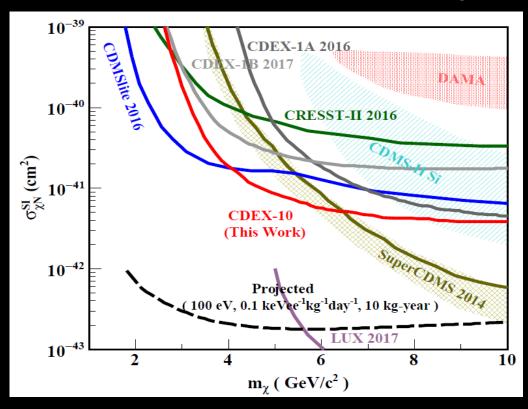


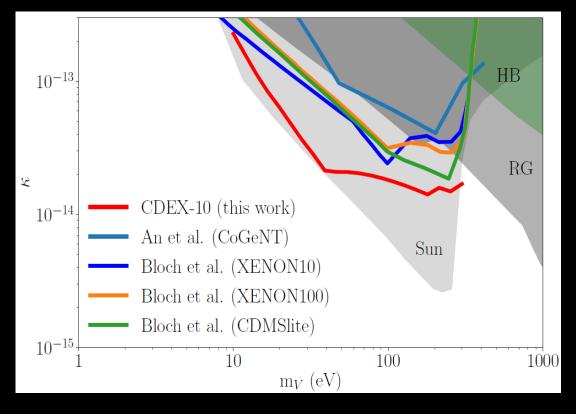
CDEX-10: ~10kg PPC Ge array



CDEX-10 results

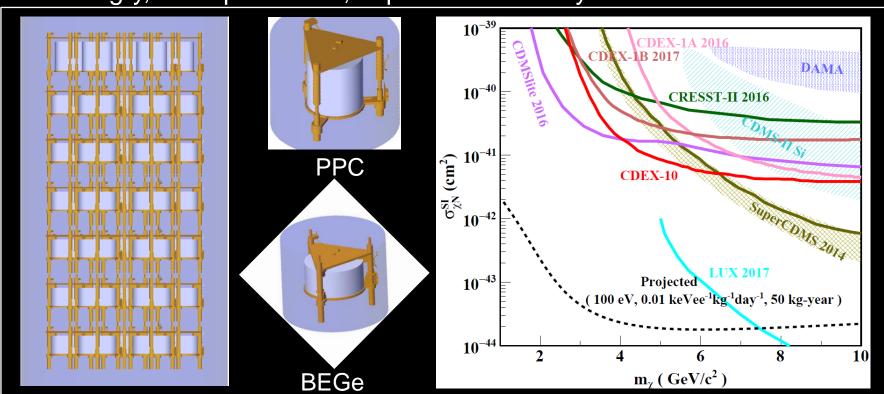
- Threshold 160 eV
- Spin-independent results: competitive in 4-5 GeV/c²
- Solar dark photon limit: improving direct detection results in 10-300 eV/c²

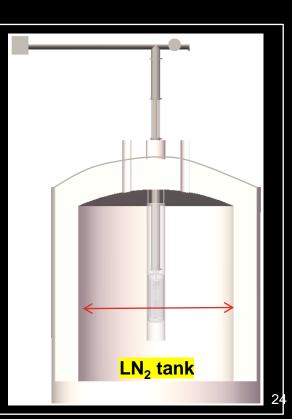




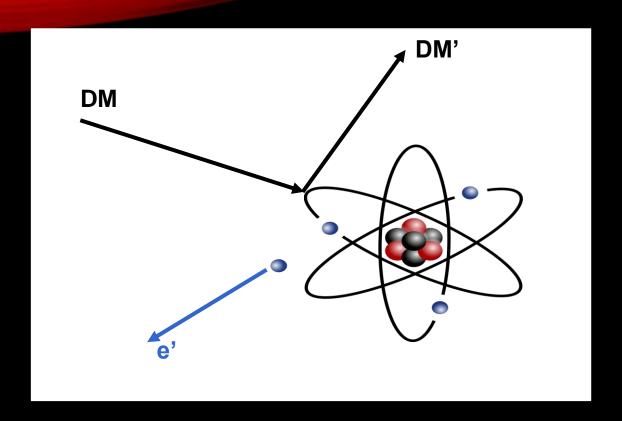
CDEX-50 Plan

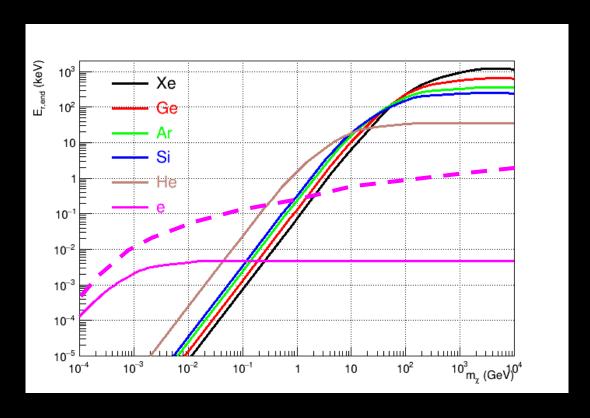
- CDEX-50 for DM search
 - Natural Ge
 - Immersed in LN
 - PPC+BEGe detectors
 - 50kg-y, 0.01cpkkd BKG, expected sensitivity ~10⁻⁴⁴cm²





Going after electrons

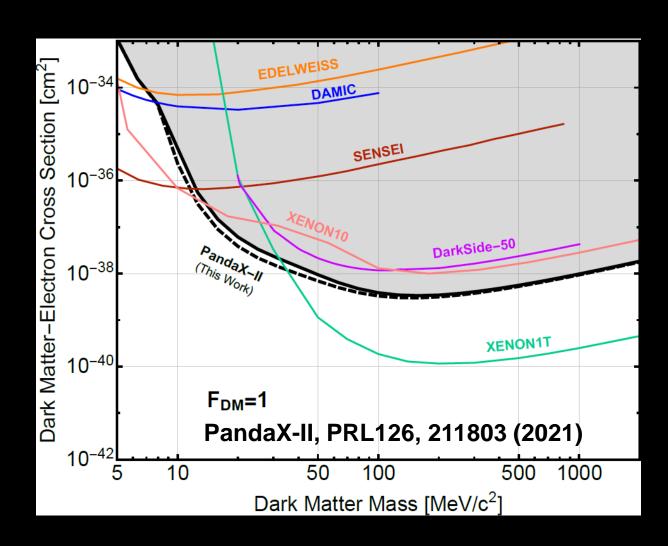




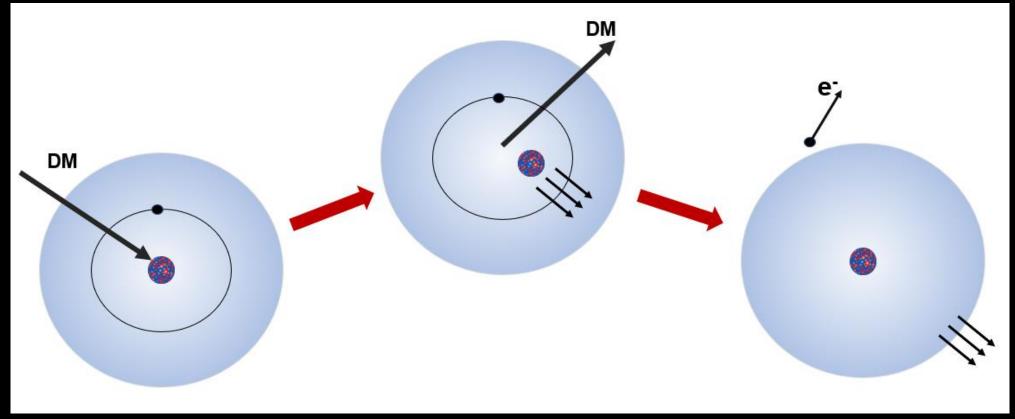
- DM could certainly interact directly with electrons
- Bound electrons in the atomic, leading to sizable energy deposition from electron ionization. Essig, Mardon, Volansky, PRD 85, 076007 (2012)

Latest search from PandaX

- Ionization-only signals (S2) in LXe TPC can see single e⁻!
- Large target mass
- Lowered threshold to 80 eV
- Best constraint to DM-e interaction between 15-30 MeV/c²

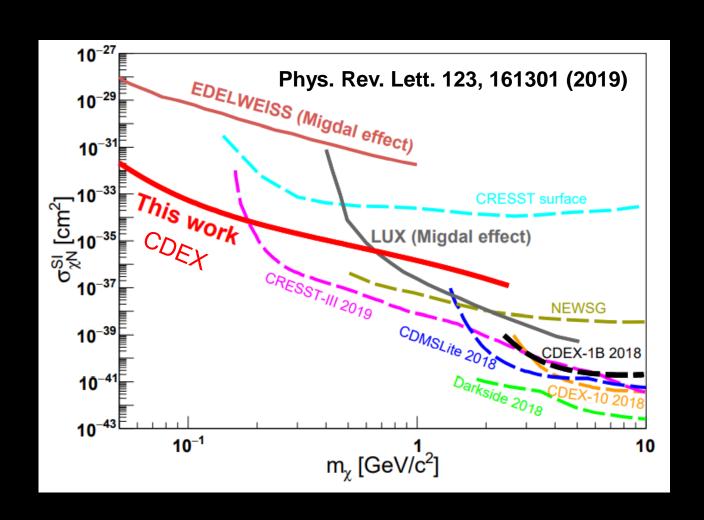


Migdal effects in DM-N scattering



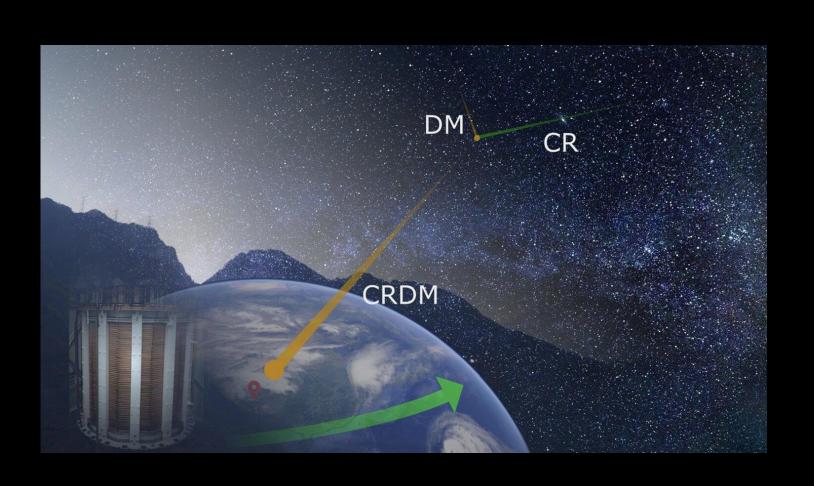
- Reformulated: Abe et al., JHEP 2018, 194 (2018); Dolan et al., PRL 121, 101801 (2018)
- Direct DM-e ionization and Midgal-induced ionization probability are closely related. Essig, Pradler, Sholapurkar, and Yu, PRL 124, 021801 (2020)

Low mass DM with Migdal



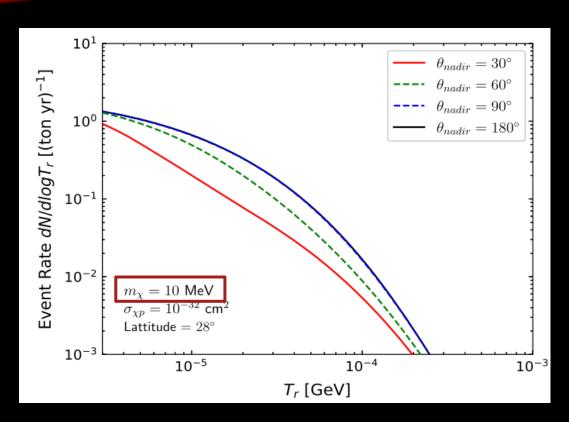
- Searches using expected signals from Migdal effect: XENON, LUX, EDELWEISS, CDEX
- Significant extension of DM-N interaction to the low mass region
- Number of experimental efforts to really see such an effect: Majewski, TAUP 2021; Nakamura et al., PTEP 1, 013C01 (2021); arXiv:2112.08514

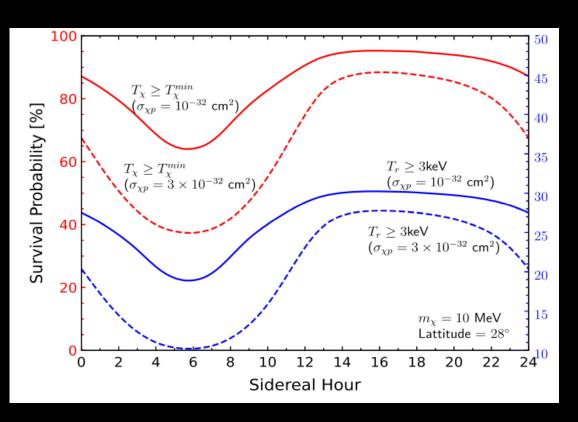
Cosmic Ray Boosted DM



- CR mostly p/He
- For finite DM-N scattering, DM acceleration by CR is inevitable!
- Bringmann and Pospelov, PRL 122, 171801 (2019)

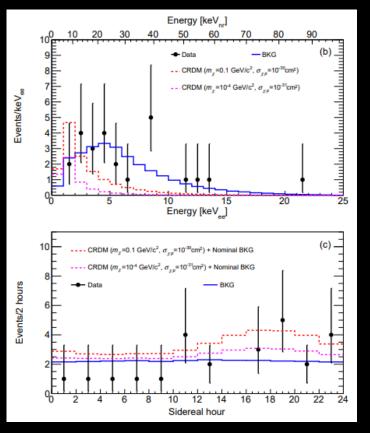
Diurnal effects

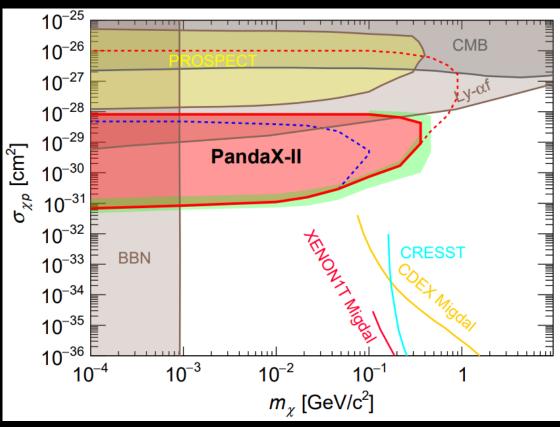




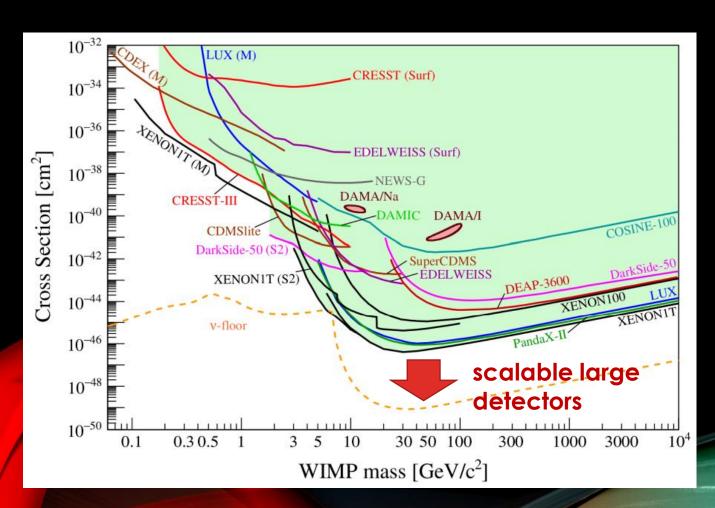
 A boost in recoil energy + a diurnal modulation in rate and E spectrum, Ge, Liu, Yuan, Zhou, PRL 126, 091804 (2021)

Search for diurnal modulation in PandaX





- First diurnal search carried by PROSPECT, PRD 104, 012009 (2021)
- PandaX: PRL 128, 171801 (2022)
- Similar search by CDEX, arXiv:2201.01704



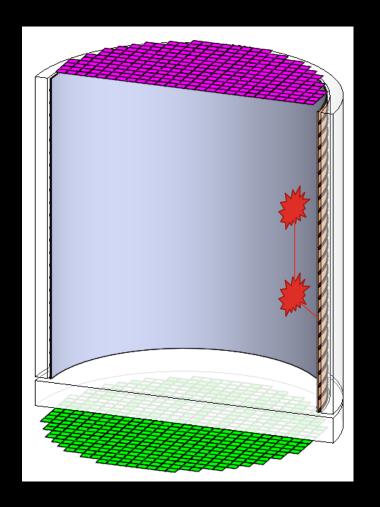
Heavy DM searches

>10 GeV or so

Heavy DM search: scalable large detectors

Material background (gammas/neutrons)

- Cluster @ boundary
- Multi-site
- Low energy scatter deep in target without high E scattering in outskirt further suppressed
- ⇒ "self-shielding" if reconstruct vertex



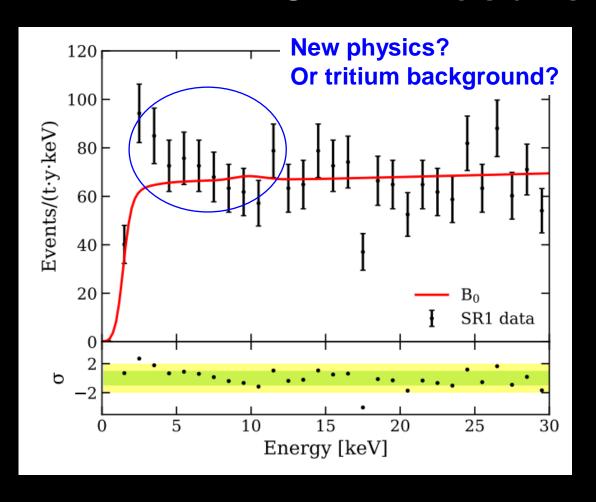
Liquid xenon experiments



Surface Neutron AC WIMP 8000 4000 2000 ${}_{2}S2_{b}$ [PE] 1000 400 200 20 10 cS1 [PE]

PRL 121, 111302 (2018) 1-tonne-year, +1σ upward fluctuation best exclusion: 4.7x10⁻⁴⁷ cm² @ 30 GeV

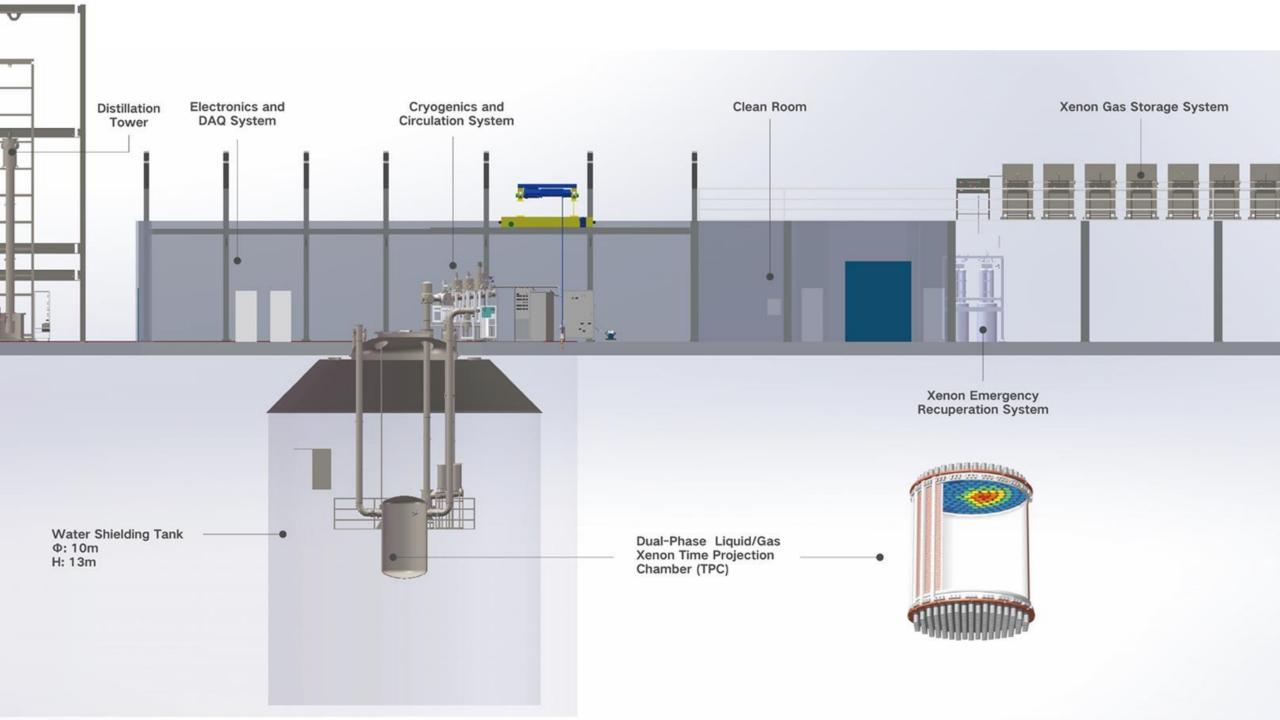
XENON1T results



PRD 102, 072004 (2020)

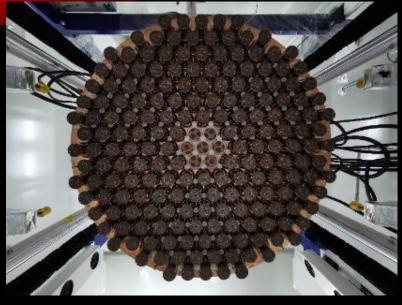
PandaX-4T milestone

- Apr., 2018, permission from CJPL management to start construction in B2 hall
- Aug., 2019, infrastructure completed, detector installation in CJPL-II started
- Mar., 2020, offline distillation of xenon completed
- May, 2020, installation completed
- Nov., 2020 Apr., 2021, commissioning run (Run 0)
- Nov., 2021 , first physics run (Run 1)





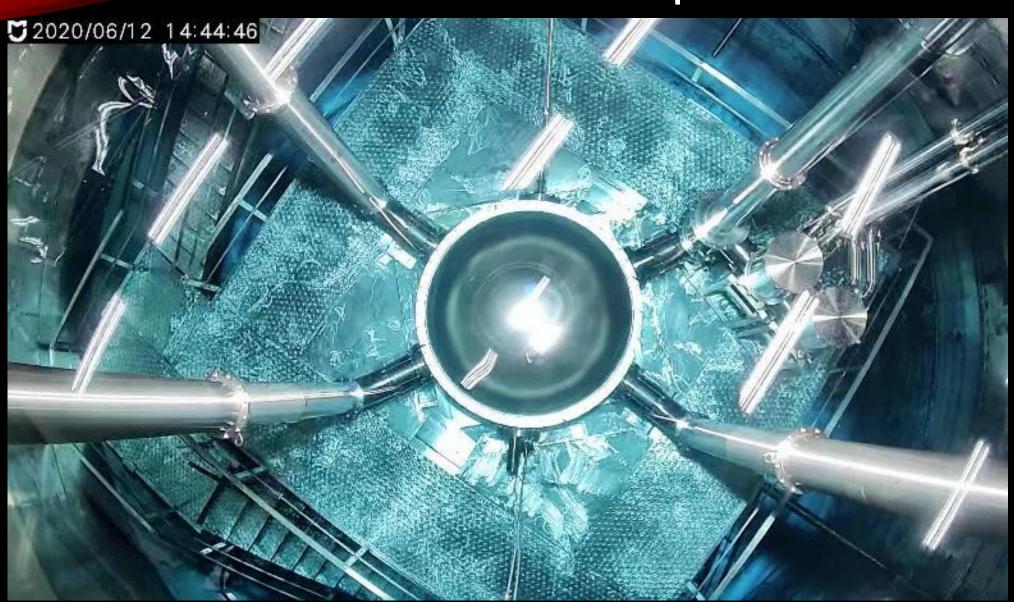
TPC installation



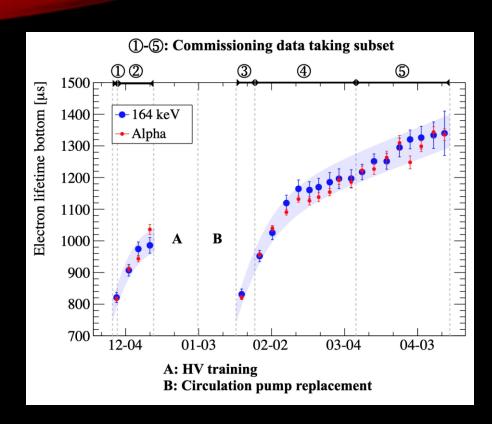




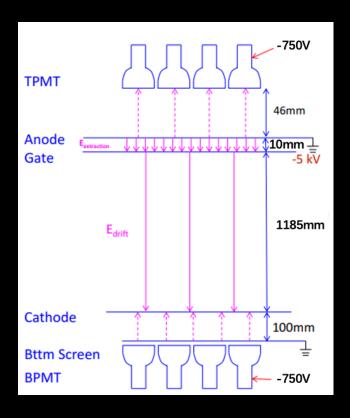
Ultrapure water filling



Run 0 configuration

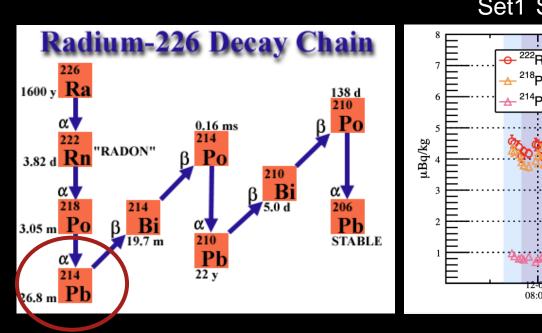


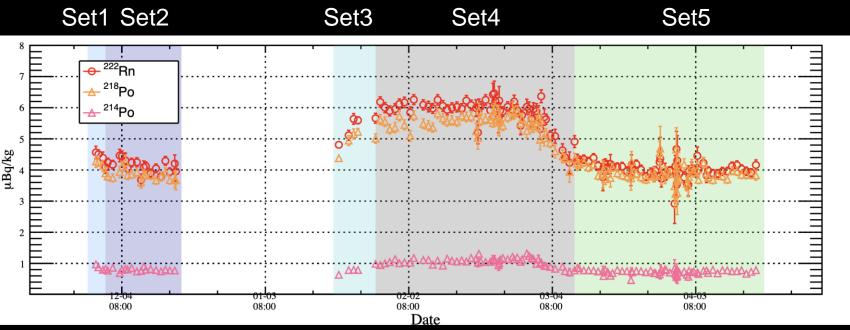
- Electron lifetime: in situ S2 vertical uniformity calibration
- Ref: the maximum drift time ~ 840 μs (field dependent)
- HV set at a few different values to avoid excessive discharges
- Stable data running period: 95.0 calendar days



	Set1	Set2	Set3	Set4	Set5
Gate(kV)	-4.9		-5	-5	
Cathode (kV)	-20	-18.6	-18	-16	

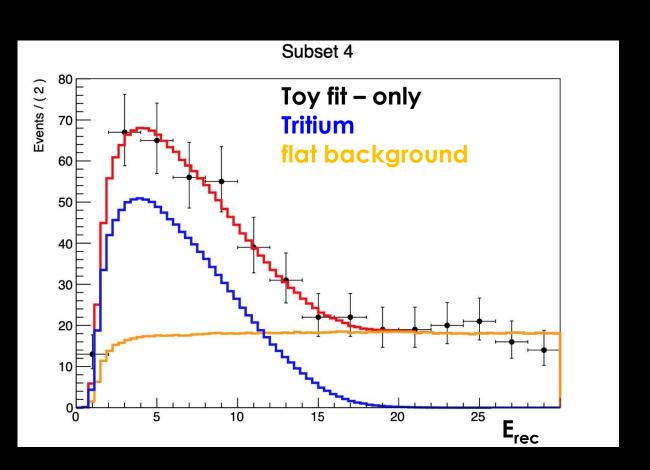
²²²Rn level evolution

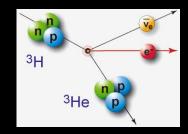




- Set 3→4: online Kr distillation (10 SLPM)
- Set 4→5: distillation off to reduce Rn emanation from the tower
- Rn level reduced from PandaX-II by 6 times

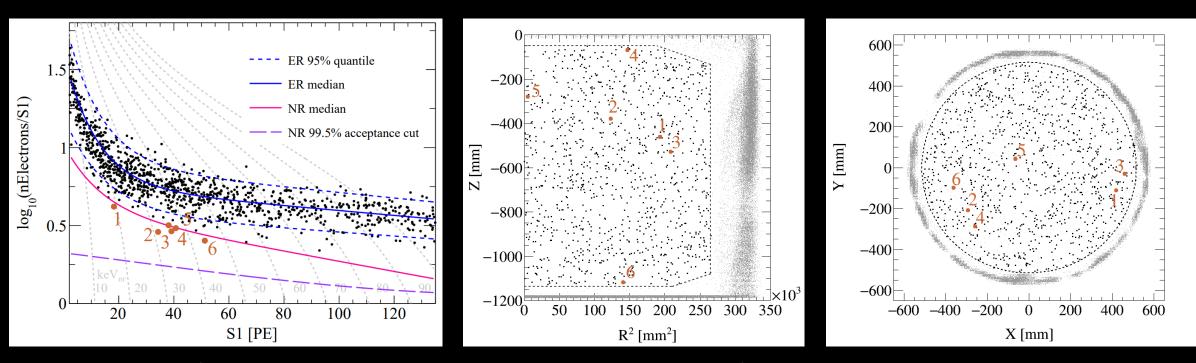
Tritium background





- Tritium spectrum identified in the data
- Likely originated from a tritium calibration at the end of PandaX-II
- Level floating in the final dark matter fit: 5(0.3)x10⁻²⁴ (mol/mol)

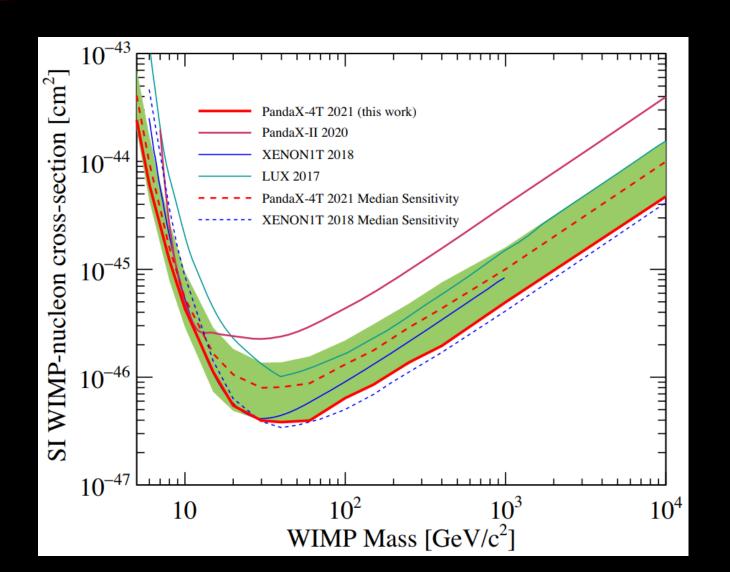
Dark matter candidates



- Events uniformly distributed in the FV, expected if dominated by tritium and radon.
- In FV, 1058 candidates, 6 below NR median line (~-1σ downward fluctuation from expected 9.8 evts)

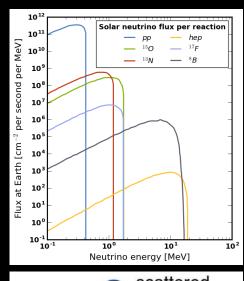
WIMP-nucleon SI exclusion limits

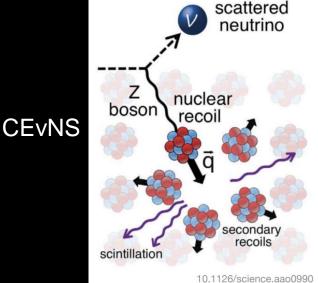
PRL 127, 261802 (2021)

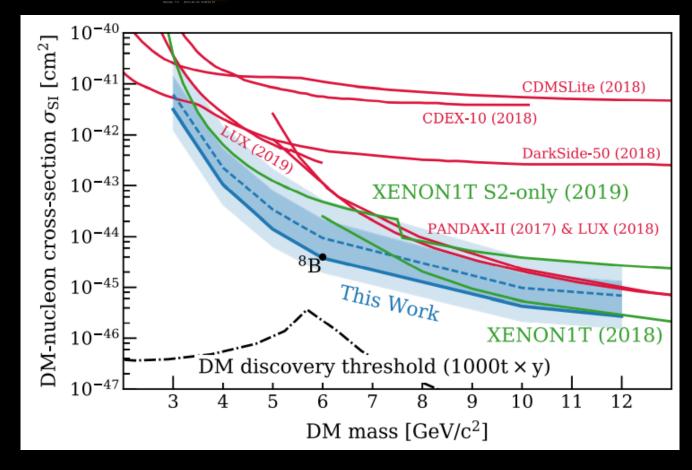


- Exposure: 0.63 tonne•year
- Sensitivity improved from PandaX-II final analysis by 2.6 times (40 GeV/c²)
- Strongest exclusion limit to date (downward background fluctuation by -1σ)

⁸B neutrino floor?



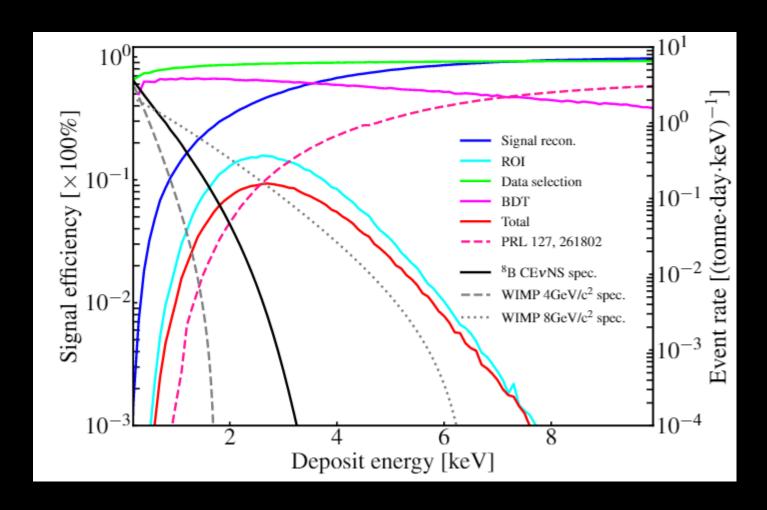




XENON1T, 1-tonne×year, PRL 126, 091301 (2021)

PandaX-4T Search on B8 CEvNS

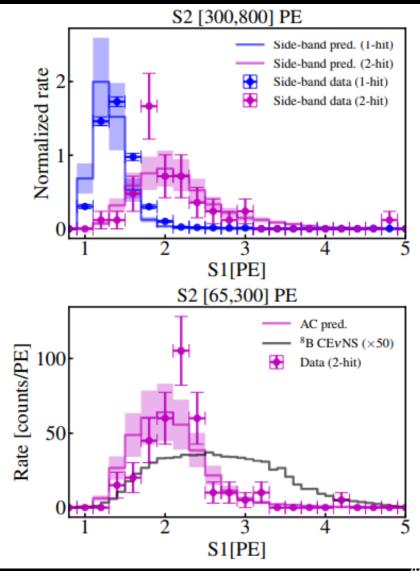
- To enhance sensitivity ⁸B, need to lower the selection threshold (S1↓, S2↓)
- Key difficulty: accidental background 1
- Blind analysis: with 0.48 ton-year data, excluding period when we see an increase of noises rates (micro-discharge)



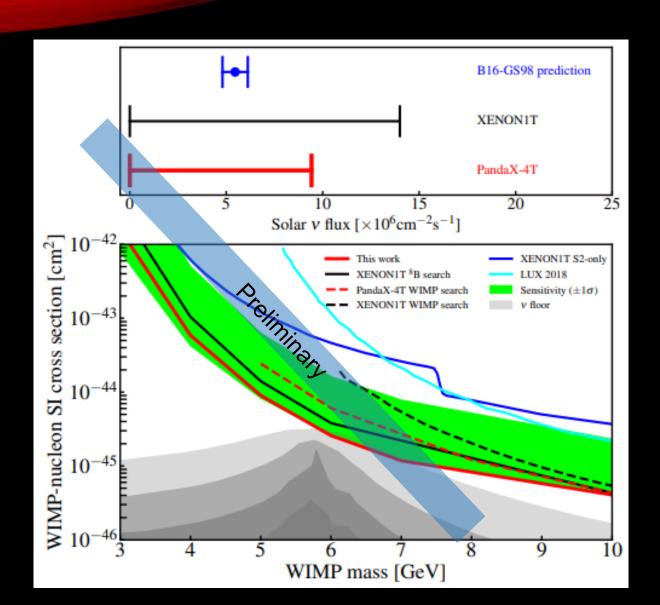
Control of accidental background

- Use "scrambled" real data to predict accidental background
- A machine-learning multivariant cut (BDT) was developed to further suppress such background from B8 signal
- Training/selection is entirely blinded
- 20% probability (with 8B CEvNS and no WIMP)

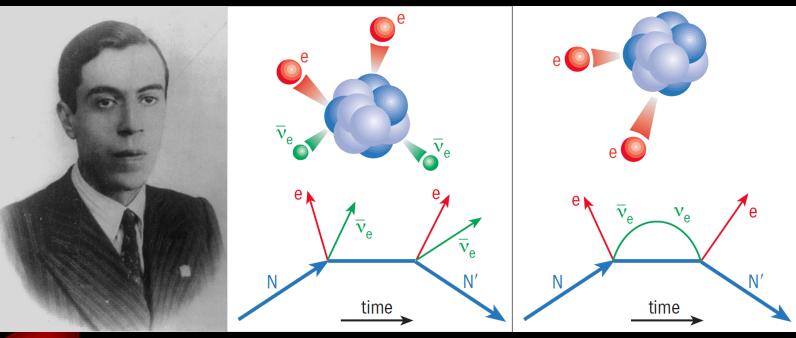
Nhit requirement on S1	BDT	Expected BKG (evt)	Expected 8B (evt)	Observed (evt)
2	Before	62.6	2.3	59
	After	1.5	1.4	1
3	Before	0.9	0.4	2
	After	0.04	0.3	0



⁸B & low mass WIMP results



- Leading constraint on ⁸B flux using CEvNS
- Can cast constraint on neutrinonucleus interactions
- Assuming nominal ⁸B background, also set tightest low mass WIMP-nucleon SI interaction limit between 5 and 10 GeV/c²
- In arXiv this week



Neutrinoless double-β decay

 $\bar{\mathbf{v}} \neq \mathbf{v}$

From Physics World

OvDBD, if found

- Majorana or Dirac
- Lepton number violation
- Measures effective Majorana mass: relate 0vββ to absolute neutrino mass

$$(T_{1/2}^{0\nu})^{-1} = G^{0\nu}(Q,Z) \ |M^{0\nu}|^2 \ \frac{|\langle m_{\beta\beta}\rangle|^2}{m_e^2}$$
 Nuclear

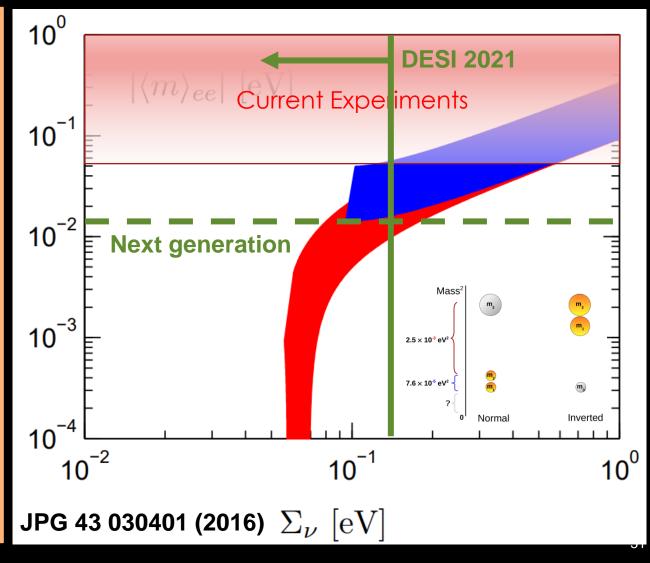
Phase space factor

Effective Majorana neutrino mass:

$$|\langle m_{etaeta}
angle|=\left|\sum_{i=1}^3 U_{ei}^2 m_i
ight|$$

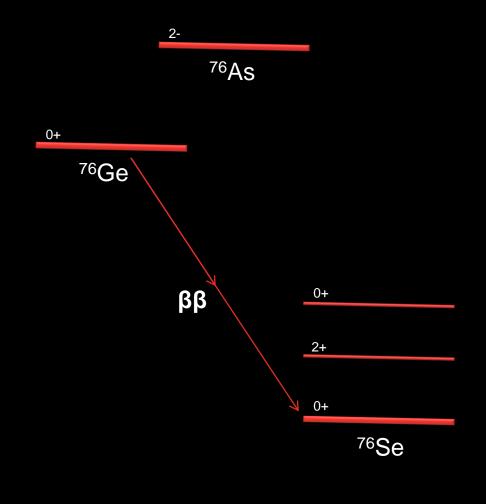
matrix

element

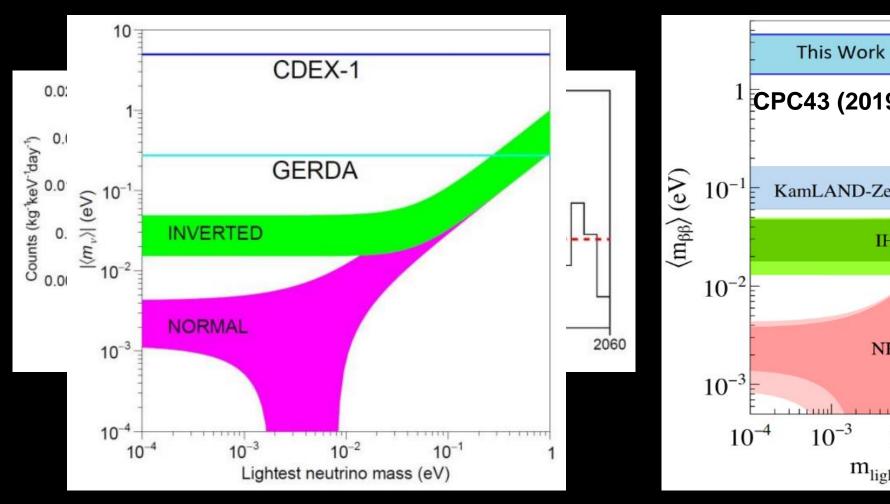


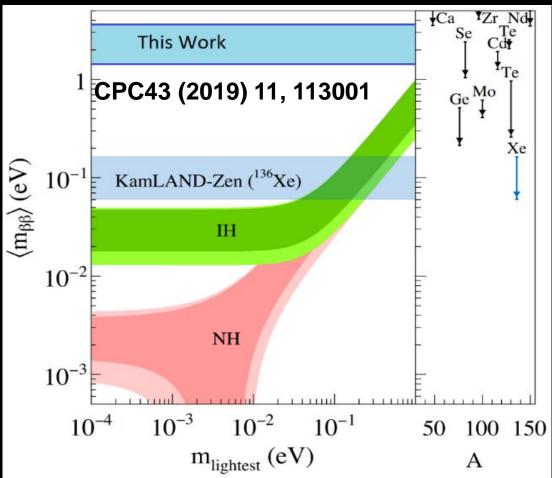
$\beta\beta$ isotopes

Isotope	Q-value [MeV]	Natural abudance [%]	
⁴⁸ Ca	4.27	0.187	
¹⁵⁰ Nd	3.37	5.6	
⁹⁶ Zr	3.35	2.8	
¹⁰⁰ Mo	3.03	9.8	
⁸² Se	3.00	8.7	
¹¹⁶ Cd	2.81	7.5	
¹³⁰ Te	2.53	34.1	
¹³⁶ Xe	2.46	8.86	
¹²⁴ Sn	2.29	5.8	
⁷⁶ Ge	2.04	7.73	
¹¹⁰ Pd	2.02	11.7	

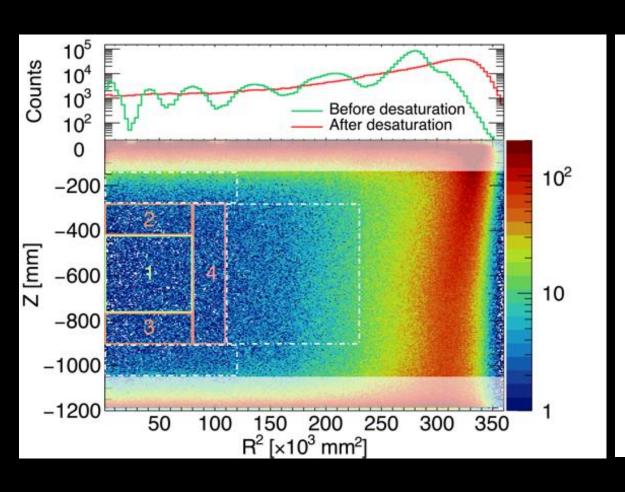


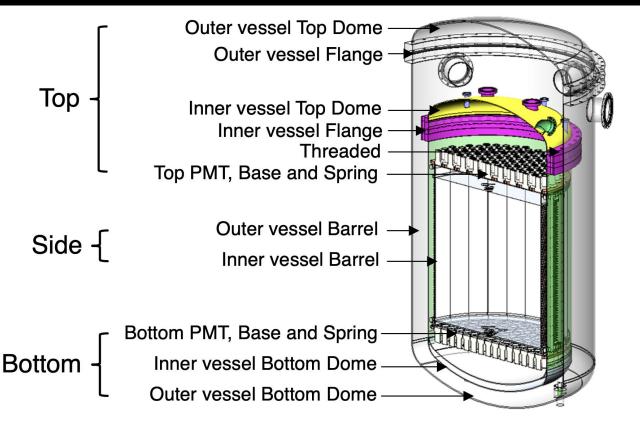
CDEX and PandaX first attempts to 0vDBD





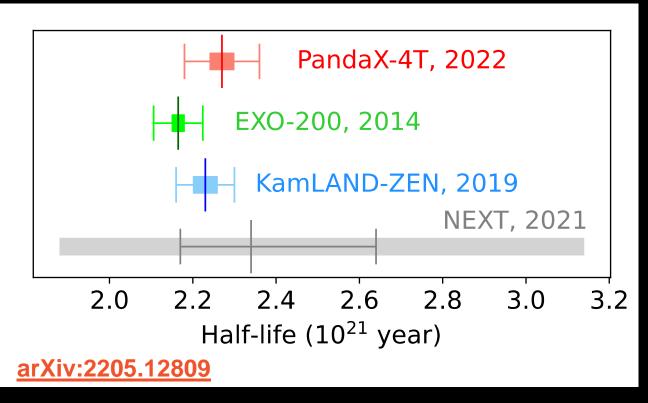
PandaX-4T, precise measurement of 2vDBD





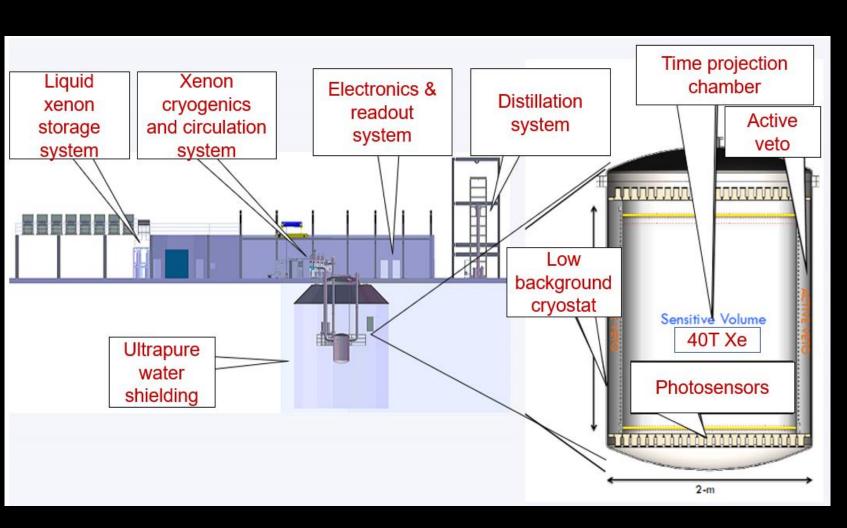
$\chi^2 / NDF = 1.10$ Best Fit ²²²Rn Counts / 10 keV Region Top ²³⁸U Bottom 40K 00 Bottom 238U Side 40K Side 238U 50 g $\chi^2 / NDF = 1.43$ Region 23 Counts / 10 keV Bottom 40K 150È Bottom 238U Side 60Co 100 Side 232Th 50 Res. [σ] 1000 1500 2000 2500 Energy [keV]

PandaX-4T, precise measurement of 2vDBD



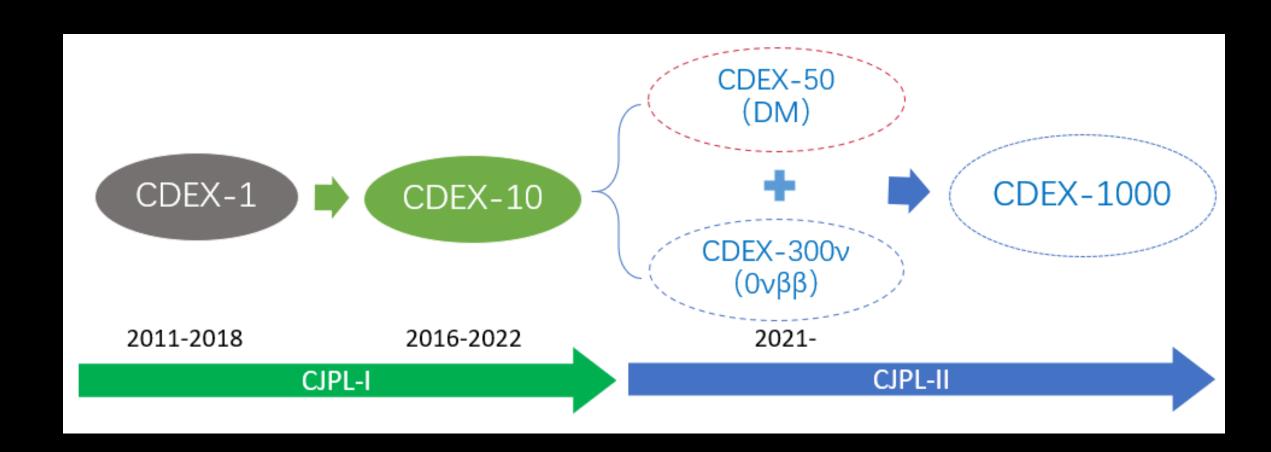
- 136Xe DBD half-life: 2.27 ± 0.03(stat.) ± 0.09(syst.) × 10²¹ year
- First such measurement with natural xenon
 - 440 keV 2800 keV range is the widest RO

PandaX next plan



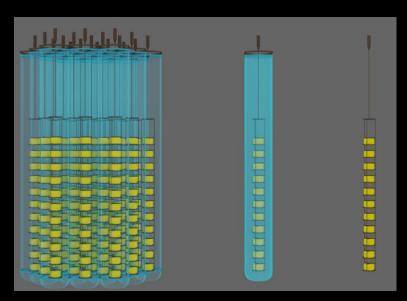
- Continue operates
 PandaX-4T till end of 2024
- Next: PandaX-xT, generalpurpose observatory on dark matter, 0vDBD (¹³⁶Xe), neutrinos, other ultra-rare phenomena
- CJPL
 - DM: unique advantage on atmospheric neutrino background (low latitude)
 - 0vDBD: Low ¹³⁷Xe background (depth)

CDEX Next Plan

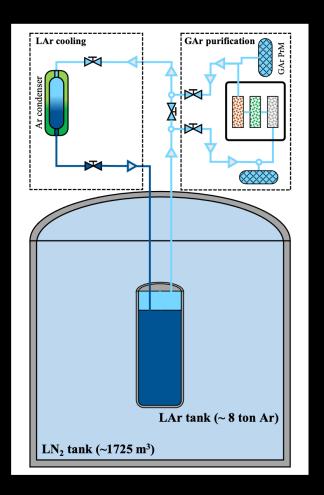


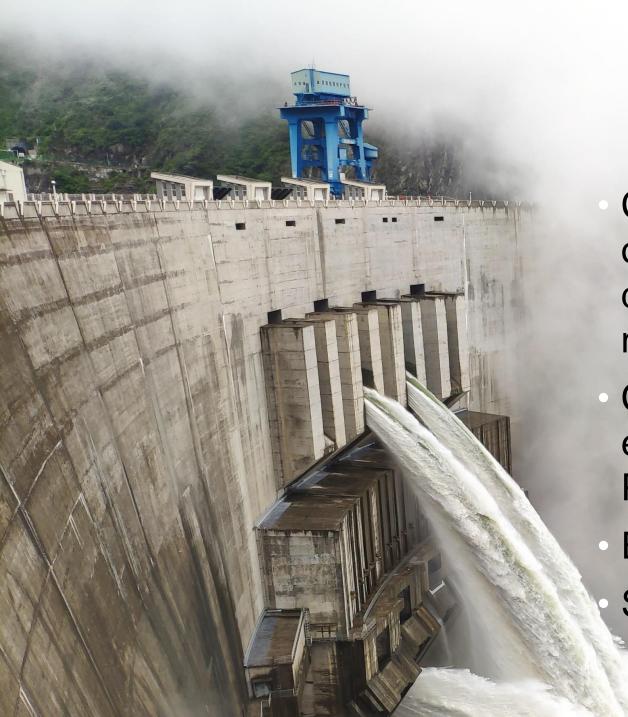
CDEX-300

- CDEX-300 for neutrinoless double beta decay
 - Ge-76 enriched >86%
 - LN + LAr (+WLS fiber + SiPM) + BEGe
 - ~200 BEGe detectors
 - Optional ICPC detector, more mass/unit
 - Data taking 2027-2030, $T_{1/2}$ > 10²⁷y, $< m_{ee} > ~50 \text{ meV}$









Summary and outlook

- CJPL has produced excellent science on the frontier of DM search, as well as demonstrating great potentials on neutrino physics
- CJPL-II upgrade is ongoing, with partial experimental occupancy in parallel (e.g. PandaX-4T)
 - Exciting future opportunities at CJPL-II Stay tuned!